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Title: Supernova Neutral Current Signal in 40Ar

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Supernova Neutral Current Signal in ⁴⁰Ar

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Abstract

We present calculations for the expected supernovae neutrino signals in a liquid argon neutrino detector. We compare these to the signals expected on carbon. In general, to extract the SN physics more than one signals is needed.

At Supernova Neutrino Energies E_v~0-50 MeV GT Transitions Dominate

$$\sigma_{v} = \frac{G_{F}^{2}g_{A}^{2}}{\pi} \left\langle f \middle\| \overrightarrow{\Sigma}(q) \middle\| i \right\rangle^{2}, \overrightarrow{\Sigma} = \overrightarrow{\sigma}\overrightarrow{\tau}$$

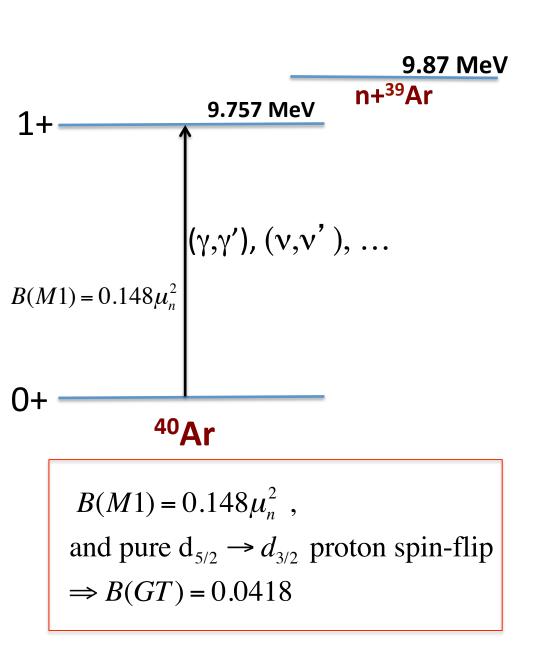
$$\xrightarrow{q=0} \frac{G_{F}^{2}g_{A}^{2}}{\pi} B(GT)$$

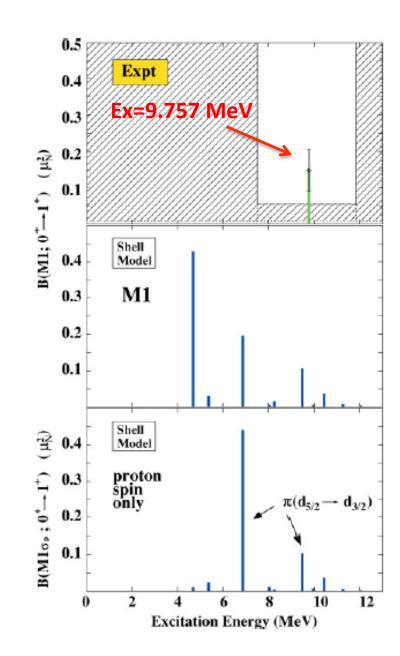
B(GT) can be approximated by the B(M1) gamma-ray strength, if the γ -transition is almost pure spin-flip

$$B(GT) = \frac{1}{2} \left(\sigma \tau \right)^{2}$$

$$B(M1) = \frac{3}{4\pi} \left\{ g_{l}^{IS} \left\langle \ell \right\rangle + \frac{g_{s}^{is}}{2} \left\langle \sigma \right\rangle + \frac{g_{s}^{IV}}{2} \left\langle \ell \tau \right\rangle + \frac{g_{s}^{IV}}{2} \left\langle \sigma \tau \right\rangle \right\}^{2} \mu_{n}^{2}$$

Almost Pure Spin-flip M1 State in ⁴⁰Ar at 9.757 MeV





1. Q-dependence of the form factors lowers cross section

2. Weak magnetism gives a ~20% difference between $v \& \overline{v}$

Four operators determine the neutral cross section to the 9.757 MeV 1⁺ in ⁴⁰Ar

$$T_{J=1M}^{mag} = \frac{q}{M_n} \Big[F_1^V \Delta_{J=1}(q) - \frac{1}{2} \mu^V \Sigma_{J=1}^{'}(q) \Big] \qquad \left\langle d_{3/2} \| \Delta \| d_{5/2} \right\rangle = \frac{1}{\sqrt{4\pi}} \sqrt{10} \frac{1}{5} \left(1 - \frac{2}{5} y \right) \exp(-y)$$

$$M_{J=1}^5 = \frac{q}{M_n} \Big[F_A \Omega_{J=1}(q) - \frac{1}{2} (F_A - \omega F_p) \Sigma_{J=1}^{"}(q) \Big] \qquad \left\langle d_{3/2} \| \Sigma' \| d_{5/2} \right\rangle = \frac{1}{\sqrt{4\pi}} \sqrt{10} \frac{4}{5} \left(1 - \frac{11}{10} y + \frac{1}{5} y \right) \exp(-y)$$

$$C_{J=1}^5 = \left[F_A - \frac{1}{2} \left(\frac{q}{M_N} \right)^2 M_N F_p \right] \Sigma_{J=1}^{"}(q) \qquad \left\langle d_{3/2} \| \Sigma'' \| d_{5/2} \right\rangle = \frac{1}{\sqrt{4\pi}} \sqrt{5} \frac{4}{5} \left(1 - \frac{9}{5} y + \frac{2}{5} y^2 \right) \exp(-y)$$

$$C_{J=1}^{el5} = F_A \Sigma_{J=1}^{'}(q) \qquad \left\langle d_{3/2} \| \Omega' \| d_{5/2} \right\rangle = \frac{1}{\sqrt{4\pi}} \sqrt{5} \left(1 - \frac{2}{5} y \right) \exp(-y)$$

$$C_{J=1}^{el5} = F_A \Sigma_{J=1}^{'}(q) \qquad \left\langle d_{3/2} \| \Omega' \| d_{5/2} \right\rangle = \frac{1}{\sqrt{4\pi}} \sqrt{5} \left(1 - \frac{2}{5} y \right) \exp(-y)$$

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$$C_{J=1}^{el5} = F_A \Sigma_{J=1}^{'}(q) \qquad \left\langle d_{3/2} \| \Omega' \| d_{5/2} \right\rangle = \frac{1}{\sqrt{4\pi}} \sqrt{5} \left(1 - \frac{2}{5} y \right) \exp(-y)$$

=> At Ev~ 25 MeV cross section drops below the simple
$$\frac{G_F^2 g_A^2}{\pi} B(GT)$$

2. Weak magnetism gives a ~20% difference between $v \& \overline{v}$

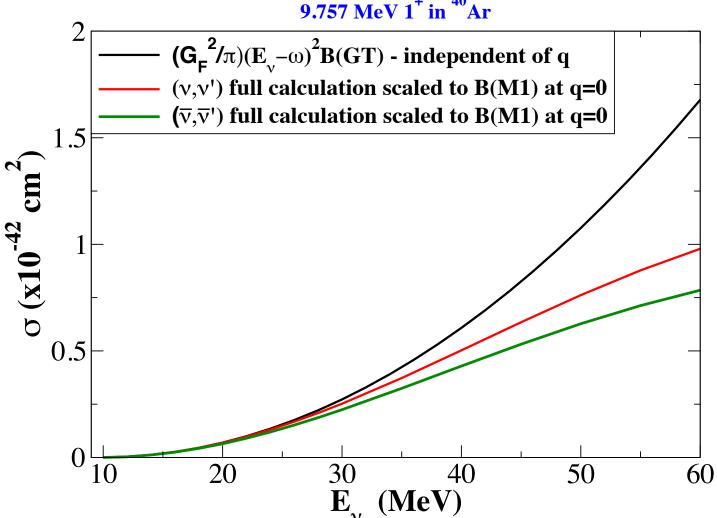
Interference term has opposite sign fo neutrino versus antineutrino

$$\sim \pm T_V^{mag}(q)T_A^{el}(q)$$

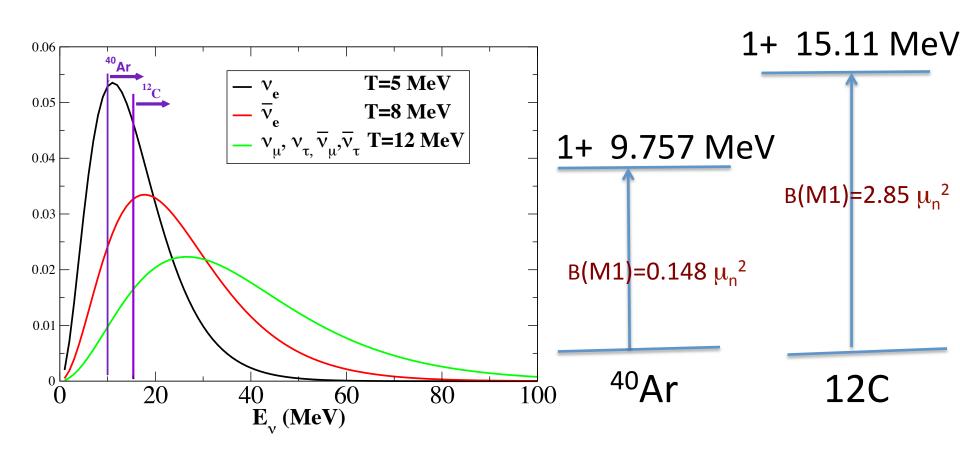
=> Neutrino and anti-neutrino cross sections deviate at Ev ~ 25 MeV

Predicted Cross Sections

Neutral current cross section 9.757 MeV 1⁺ in ⁴⁰Ar



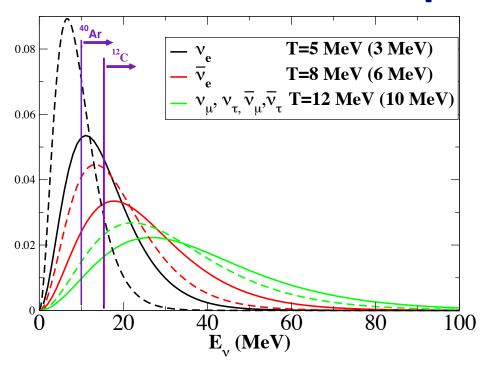
Fermi-Dirac Neutrino Flux - No Oscillations



Total Cross section (all 6 neutrinos flavors)

40
Ar $1.1x10^{-42}$ cm² < Ev> =44 MeV
 12 C $16.0x10^{-42}$ cm² < Ev> =48.8 MeV

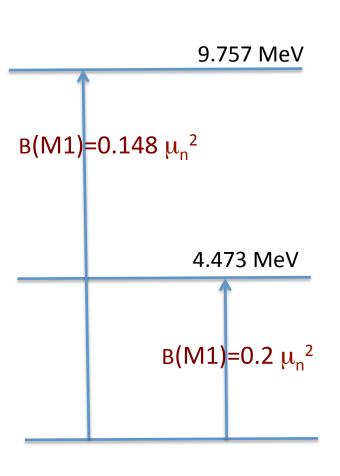
SN Neutrino Temperature Uncertain



ν_a 3-15% - also measured via CC

	σ ⁴⁰ Ar X10 ⁻⁴² cm ²		<ev> MeV</ev>		σ ¹² C	σ ¹² C X10 ⁻⁴² cm ²		<ev> MeV</ev>	
	T>	T<	T>	T<	T>	T<	T>	T<	
${ m v_e}$.07	.01	30	22	.6	0 .05	35	26.	
$ m u_{e ext{-bar}}$.18	.1	41	34	2.4	1.0	46	38	
$ u_{\mu, au ext{-bar}}$.37	.34	53	49	5.8	4.1	57	52	
$ u_{\mu, au}$.45	.3	52	48	7.3	5.0	58.	53	

Additional Neutral Current Signal in 40Ar at 4.473 MeV



9.757 MeV state almost pure $d5/2 \rightarrow d3/2$

=>(v,v) cross section straightforward

4.473 MeV structure unknown

- Needs more work
- Emitted gamma –ray could confuse (v_e ,e-) signal

To be done

- Neutral current signals of different SN temperatures, etc.
- Charged current cross sections (including qdependent form factors) and signatures
- Breakup of nucleus with neutron emission
- Analysis of signals in different detectors, water Cherenkov, etc. to max extraction of SN physics